

Extreme-scaling on Omni-Path fabric: *performance for computational astrochemistry*

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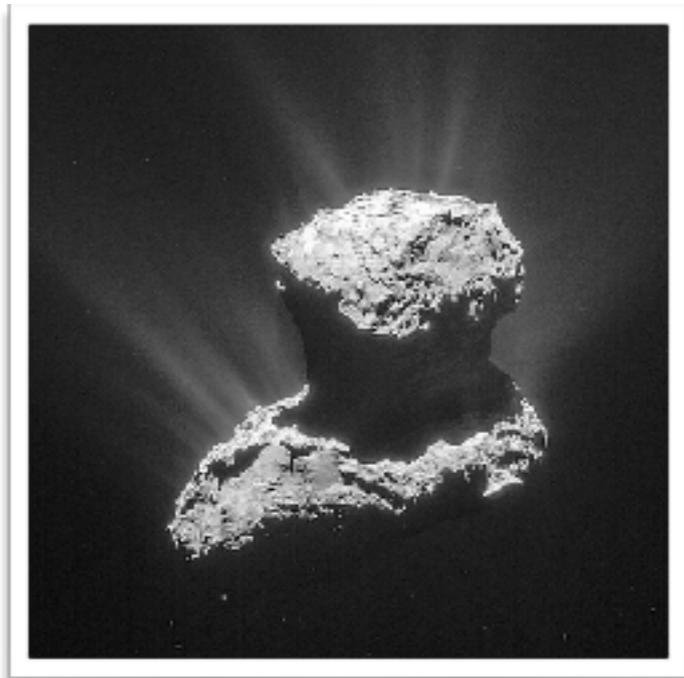
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for Astrophysics

VIPER
Transforming Research

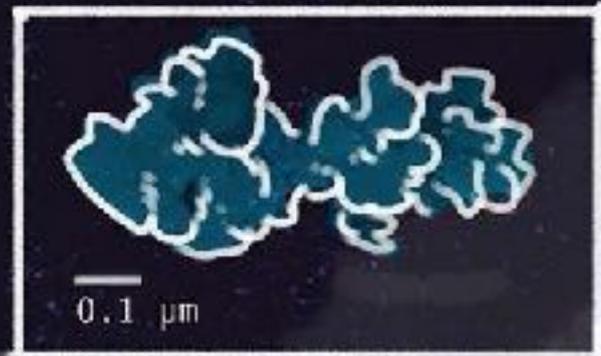
Landing on a comet – ESA Rosetta Mission

- ESA space probe launched in 2004
- Lander module Philae explored comet 67/P Churyumov-Gerasimenko in 2014



WHAT IS A COMET MADE OF?

INTERPLANETARY DUST PARTICLE



FROZEN GAS
CO
CO₂
H₂O
NH₃
CH₃OH
CH₂OH
HCN
Many More

INTERPLANETARY DUST PARTICLE

DIAMOND
GRAPHITE
SILICON CARBIDE
TITANIUM CARBIDE (TiC)
SILICON NITRIDE (Si₃N₄)
CORUNDUM (Al₂O₃)
SPINEL (MgAl₂O₄)
HIBONITE ((Ca,Ce)(Al,Ti,Mg)₁₂O₁₉)
TITANIUM OXIDE (TiO₂)
SILICATE MINERALS (OLIVINE AND PYROXENE)

COMETOPAUSE

MAGNETIC BARRIER

BOWSHOCK

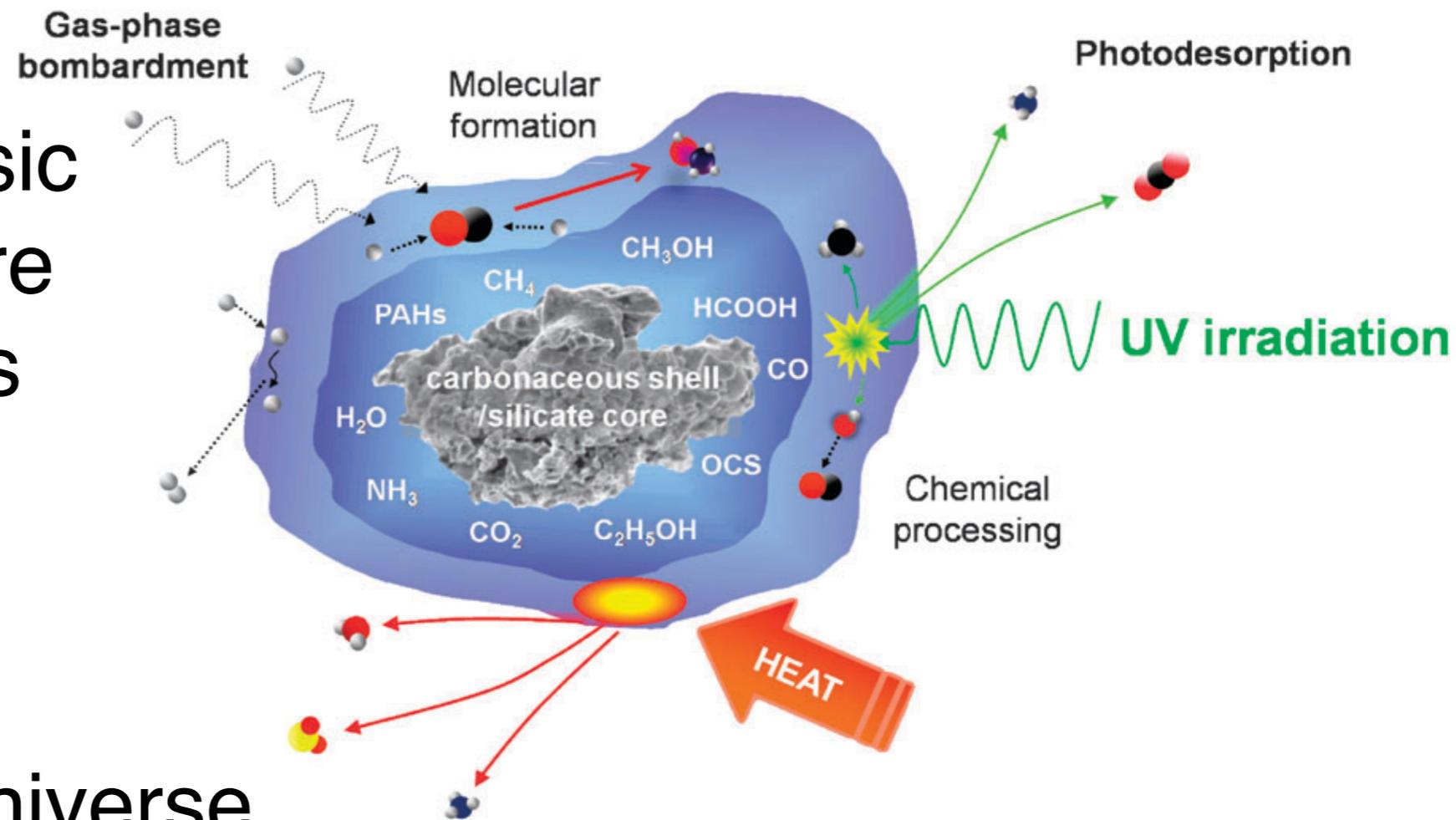
Surface chemistry on dust grains

- Dust grains acts as cosmic “**bench tops**” for chemical reactions

- Help transform basic chemicals into more **complex** molecules

- Their study is key to **understanding** low-temperature chemistry in the Universe

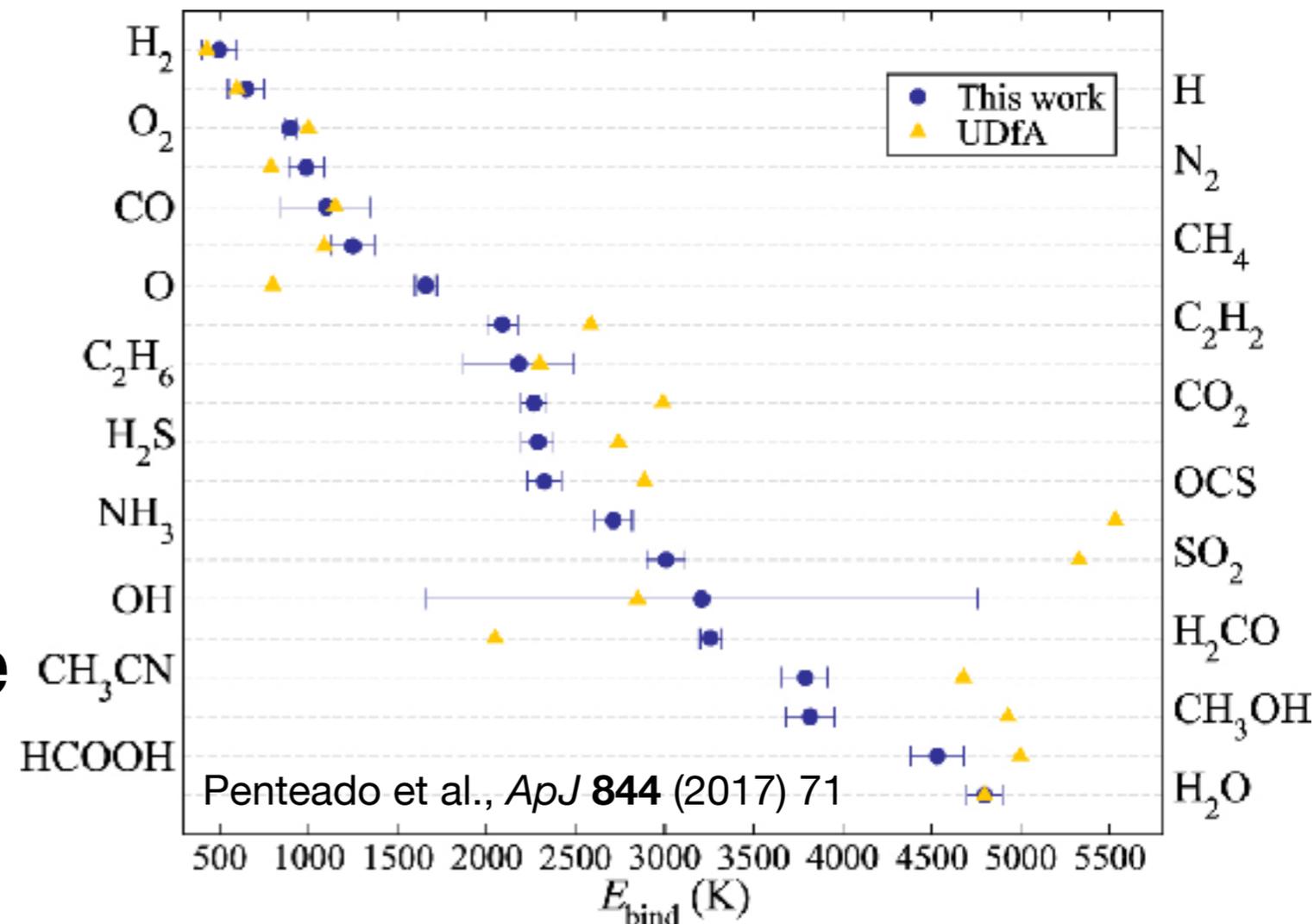
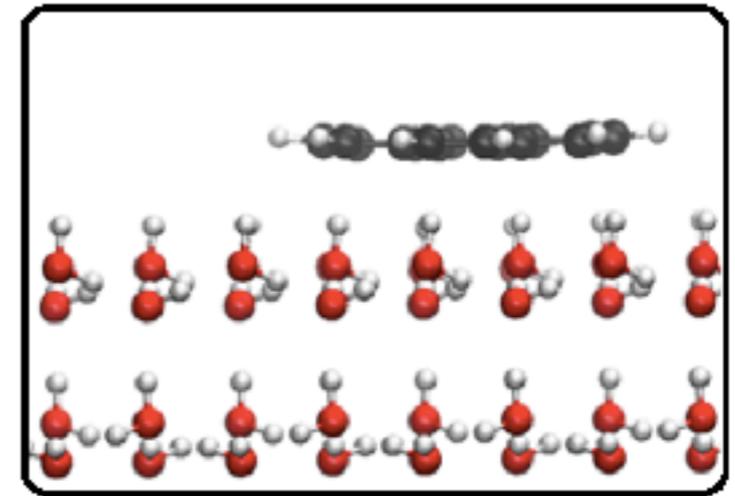
- But **how strongly** do molecules stick to ice-covered grains/comets?

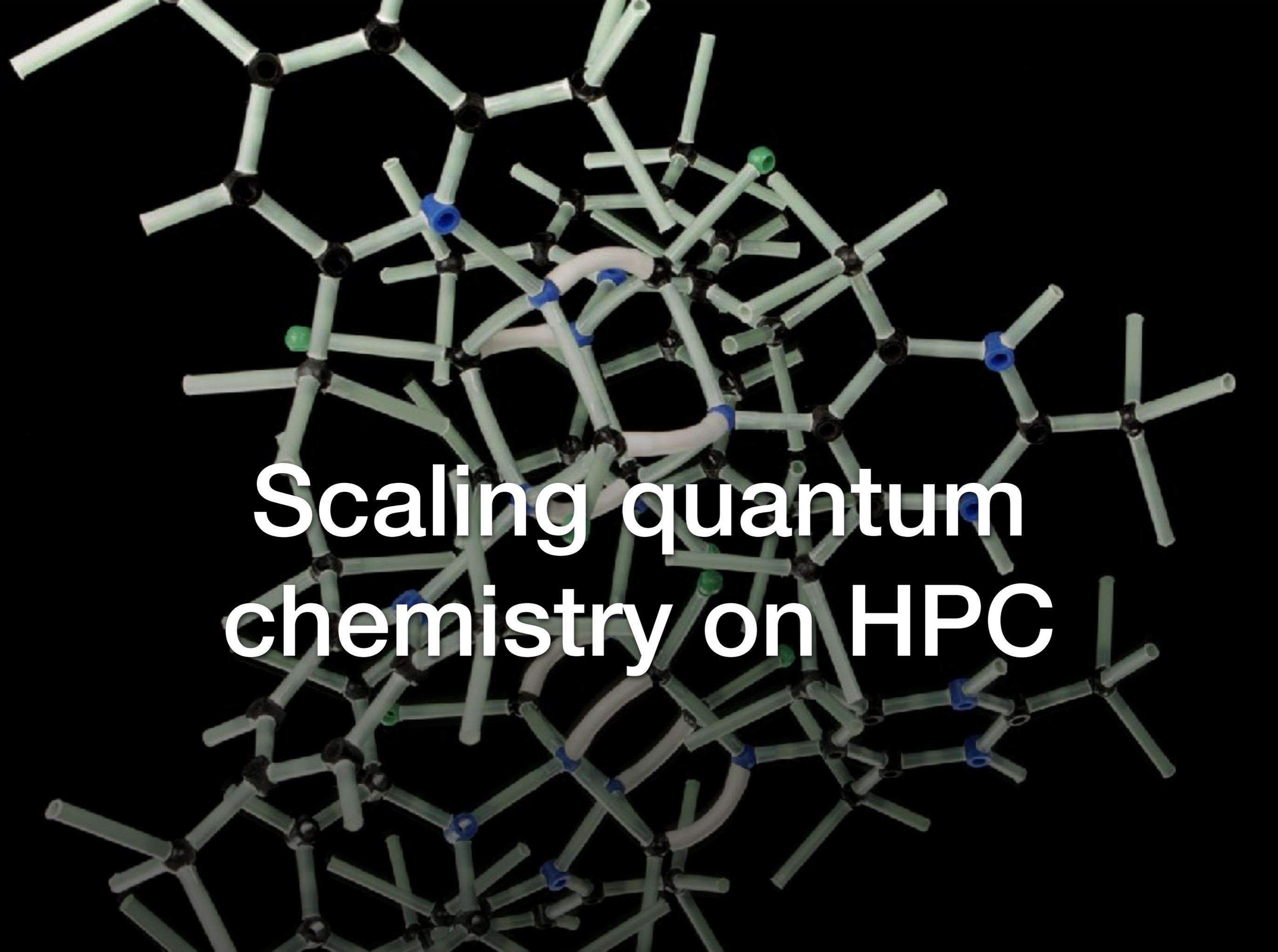


From: D.J. Burke et al., PCCP **12** (2010) 5947

Adsorption / sticking energies?

- Adsorption energies are **hard** to measure accurately
- Experimental **database estimates** are from '90s with little experimental/theoretical validation
- Quantum chemical calculations can provide a **reliable** estimate





Scaling quantum chemistry on HPC

VIPER technical profile

- 5040 Intel Broadwell E5-2860v4 (2.4 GHz) cores in 180 compute nodes
- Intel X16 100Gb/s Omni-Path interconnect
- 4 x 1 TB high memory nodes
- 2 x visualisation nodes (2 x Nvidia GeForce GTX 980 Ti per node)
- 4 x accelerator nodes (4 x Nvidia Tesla K40M GPUs per node)
- 500 TB of user storage running BeeGFS
- Each node runs a Docker container

RED OAK CONSULTING



dustervision

BeeGFS™
created by Flairhcf

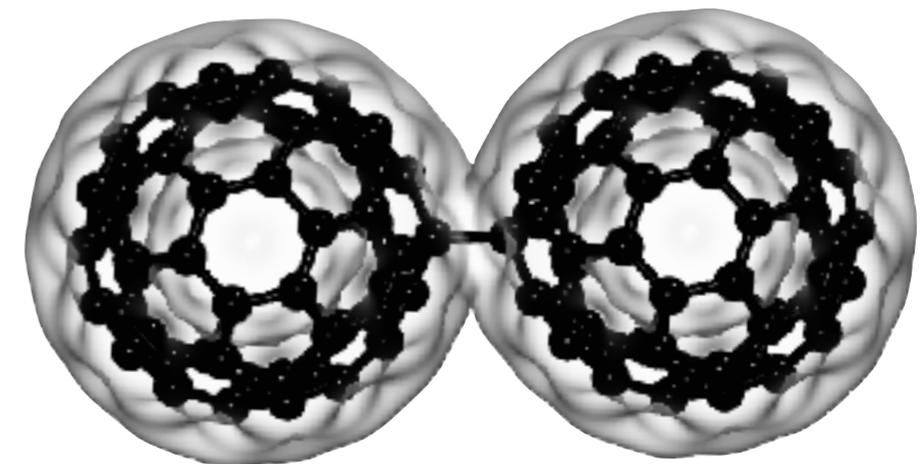
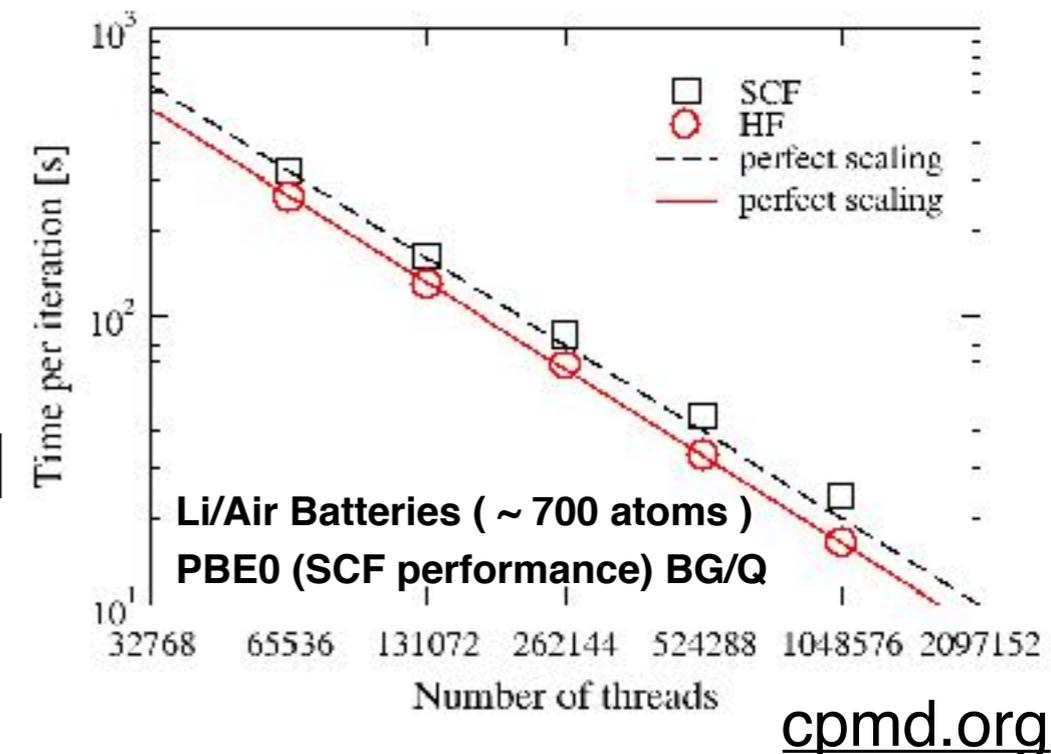
TESSERACT technical profile (SGI 8600)

- 20,256 Intel Skylake silver 4116 (2.1 GHz) cores in 844 compute nodes
- 96 GB NUMA per node (48GB/processor)
- Intel X16 100Gb/s Omni-Path interconnect
- 3 PB of Lustre file system
- DiRAC's Extreme Scaling facility
- Hosted at EPPC/Edinburgh



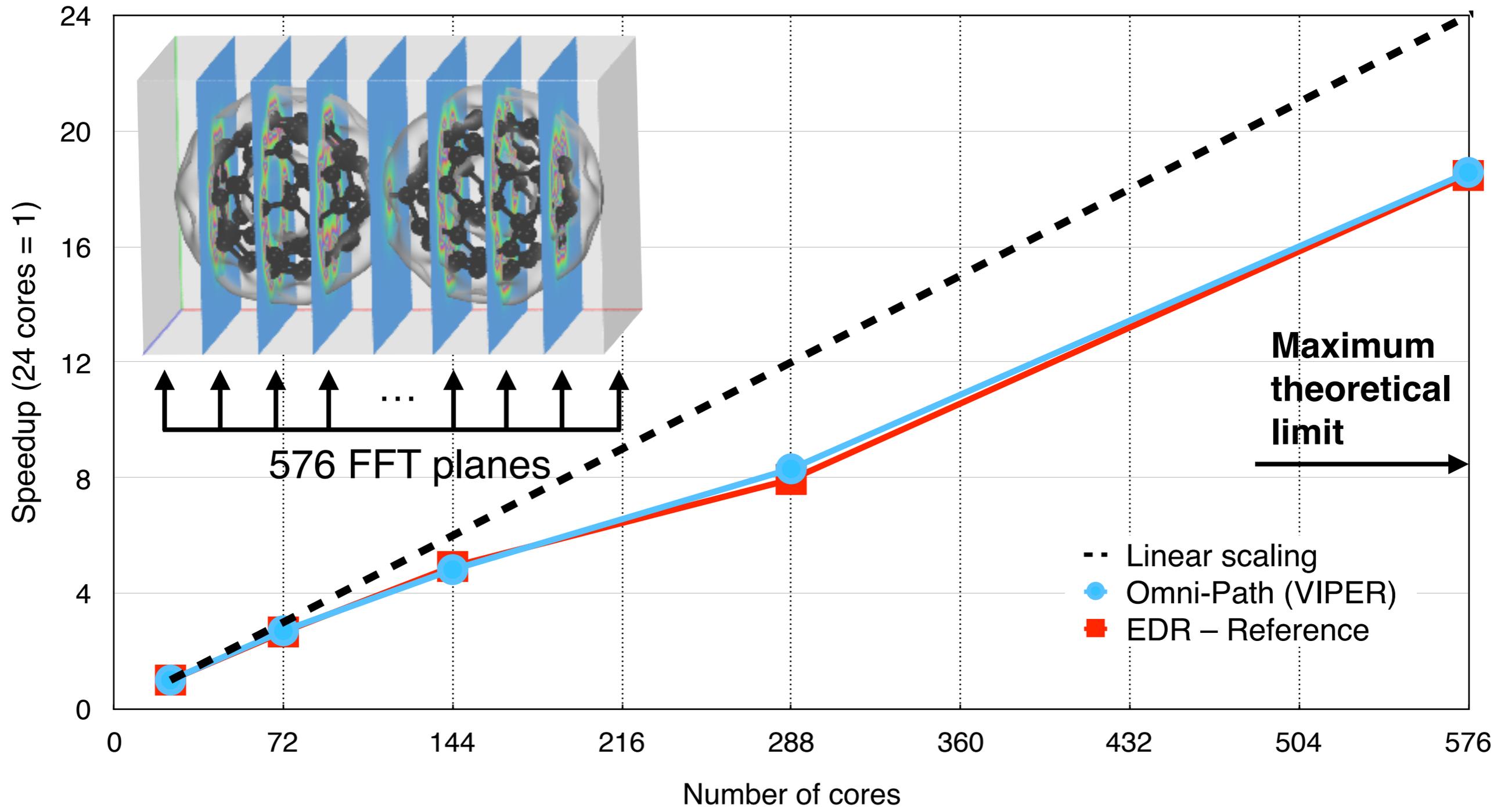
Application: CPMD – Large-scale electronic structure code

- Well established, freely available, density functional theory (DFT) code
- Developed by MPI-FKF Stuttgart&IBM
- Demonstrated scaling on large HPC
- Hybrid MPI/OpenMP parallelisation
- Bottlenecks: FFT, memory footprint and all-to-all comms
- Our system: fullerene dimer (C_{120}), 480 electrons, triplet state, B3LYP functional and a 500 Ry cutoff



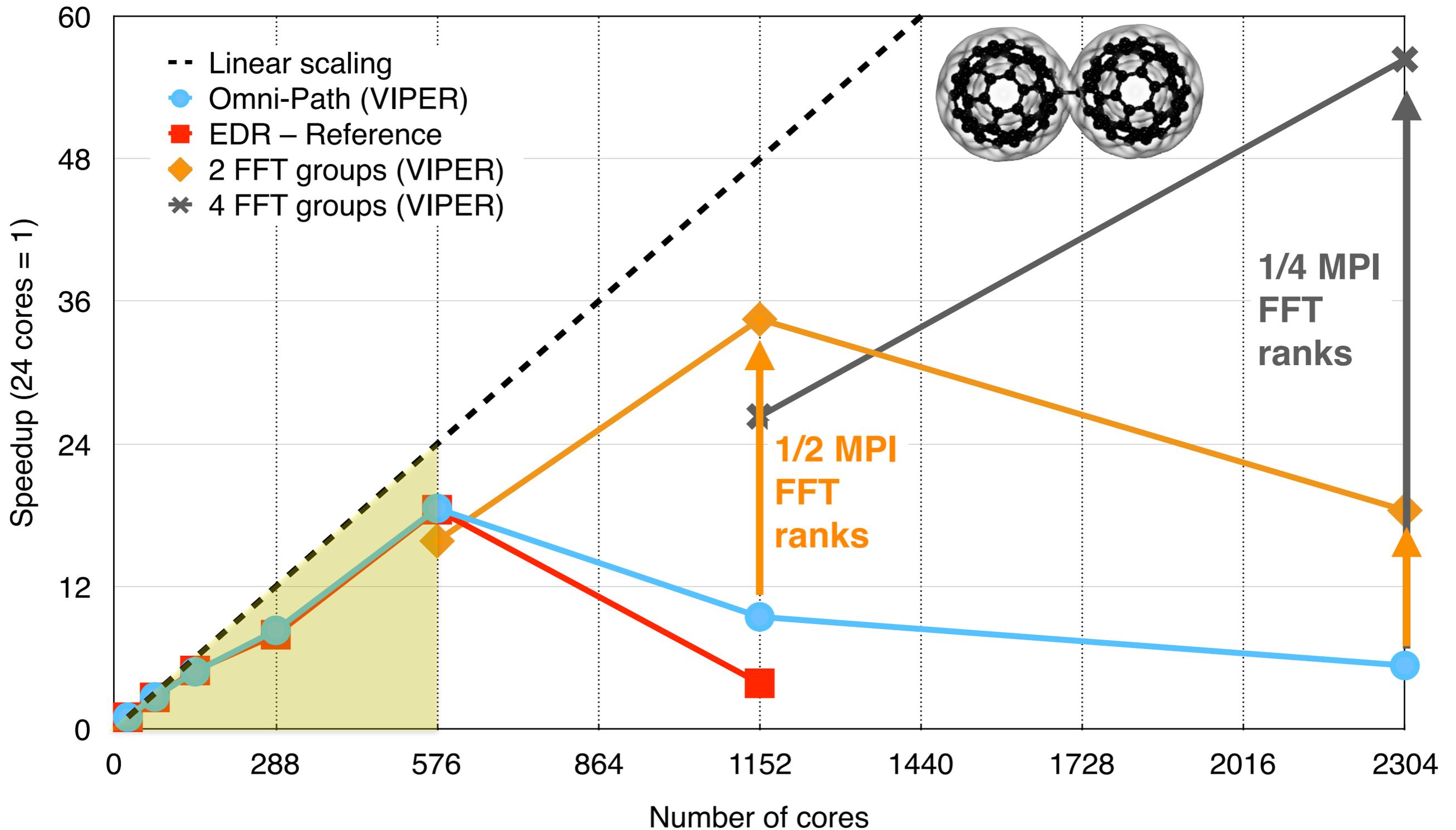
Fullerene dimer, C_{120}

Large-scale density functional theory on OPA

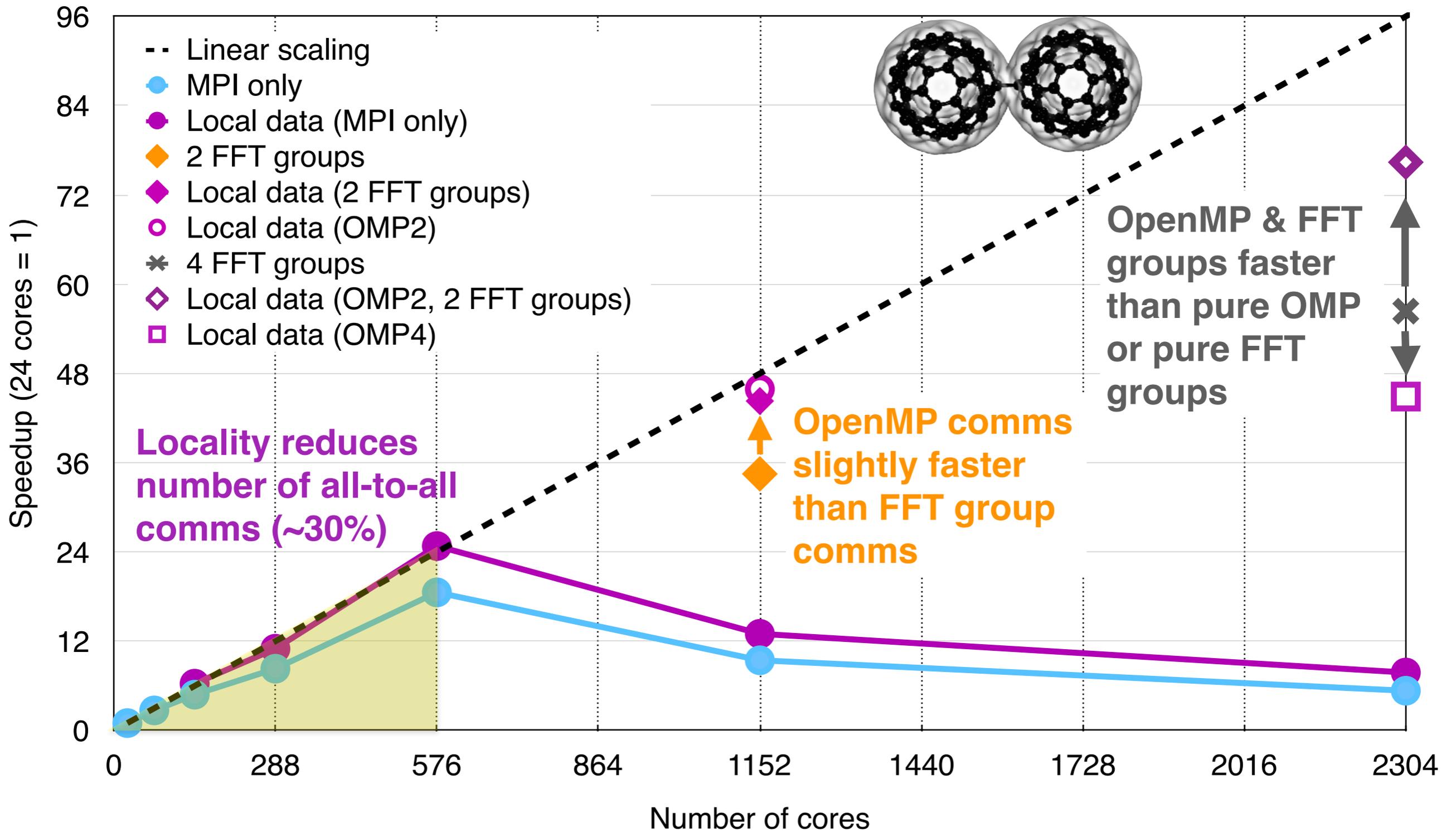


Conditions: Omni-Path HFI Silicon 100 Series; Intel 100 Series 48 port unmanaged switches; CentOS 7.2.1511; dockerised nodes; OPA 10.4.1.0-1; CPMD V4.5 compiled with ifort 2017 (-O2 -ipo -xHOST) ATLAS (single thread) and scaLAPACK

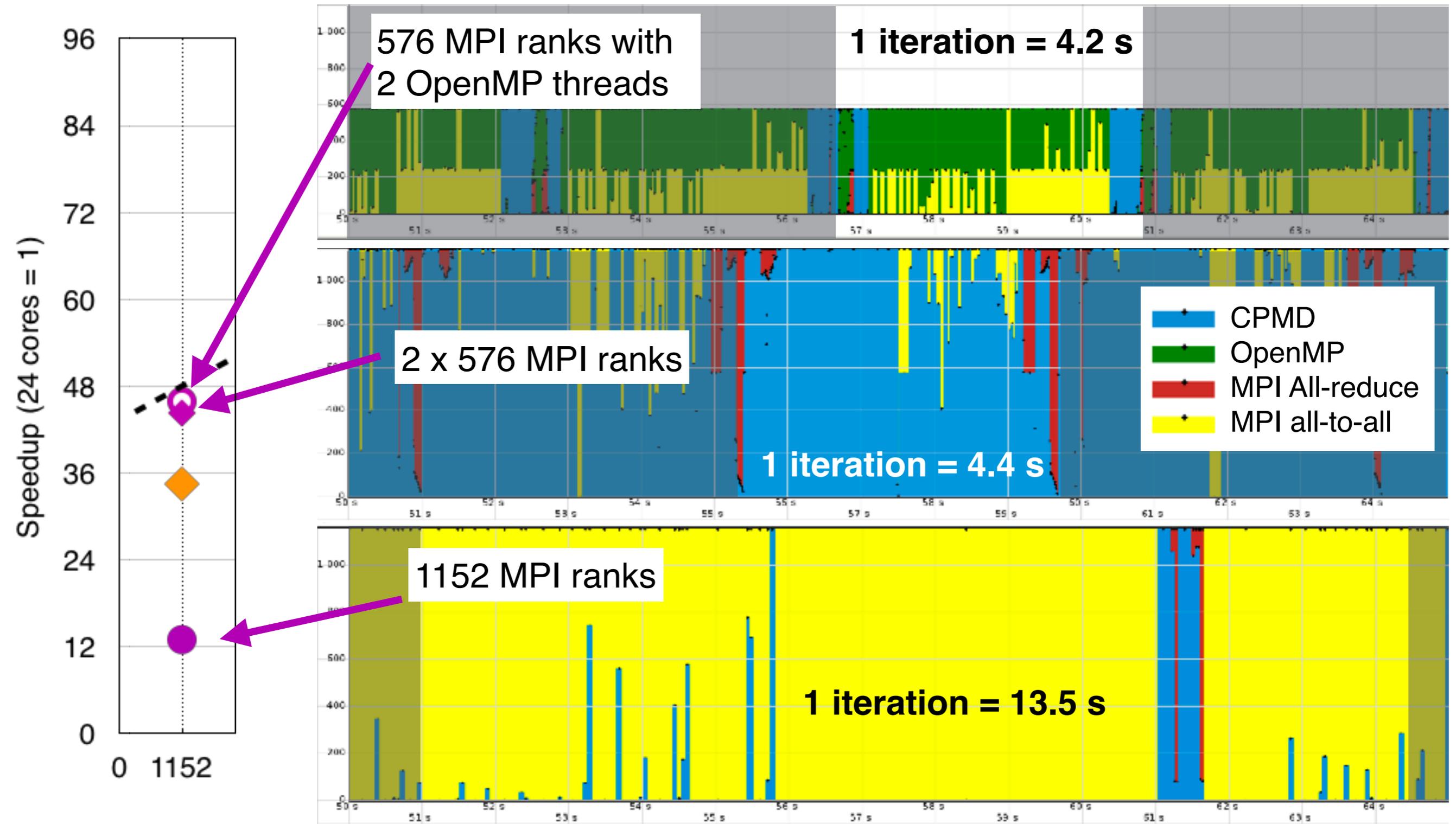
Scaling past the maximum theoretical limit



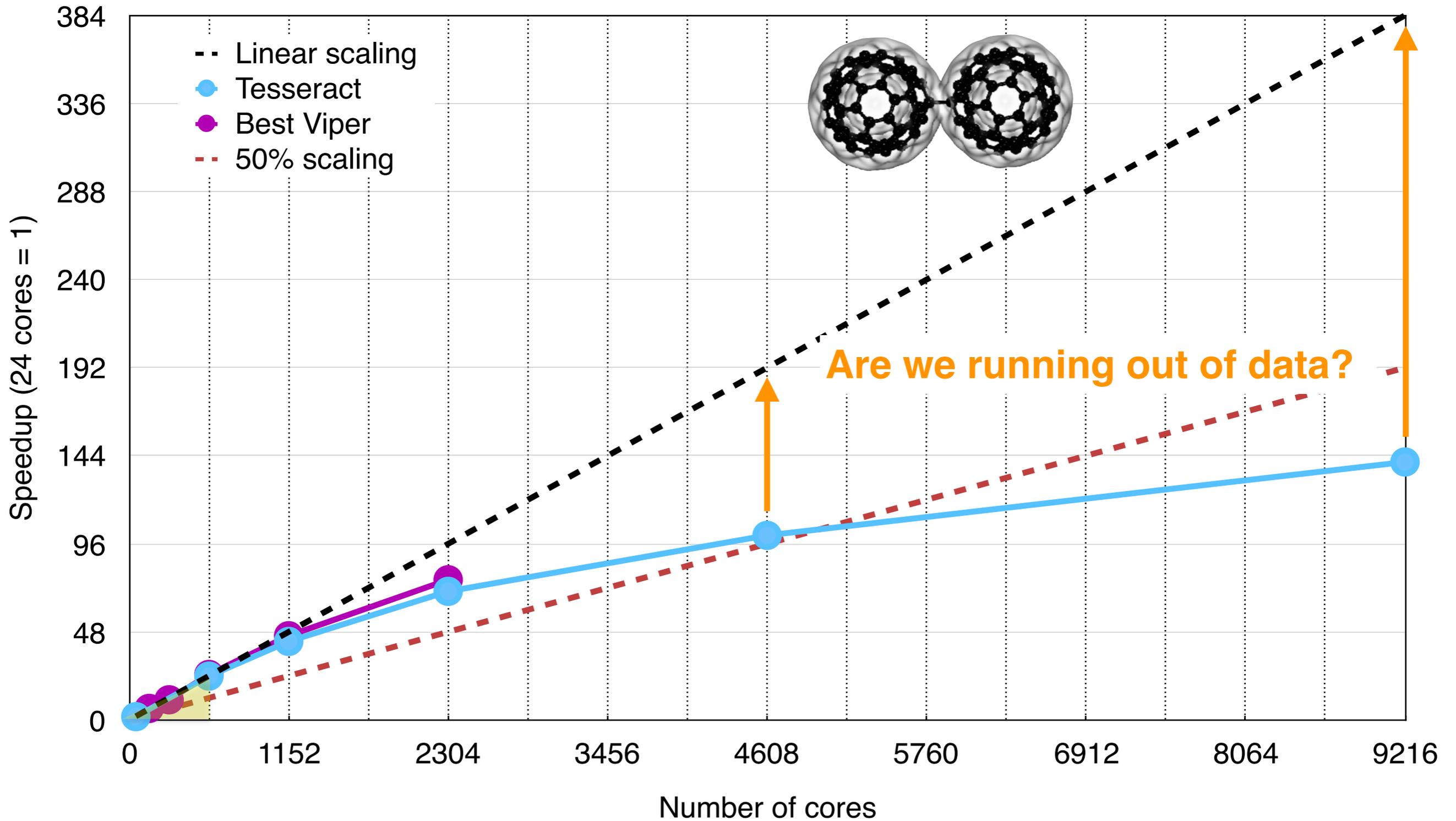
Locality and OpenMP (kitchen sink approach...)

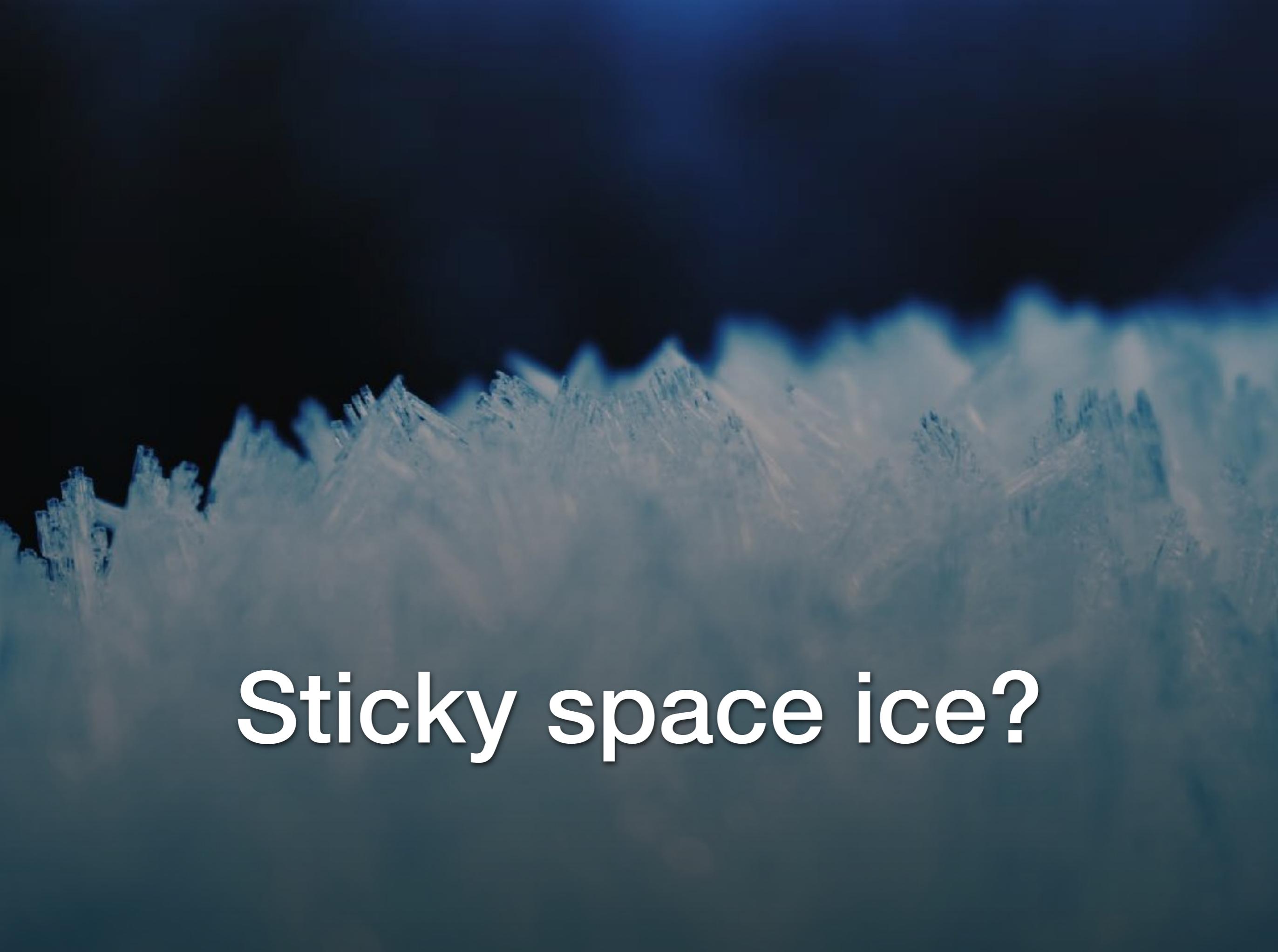


ITAC MPI trace – What happens at 1152 cores?



More cores better? – Tesseract scaling

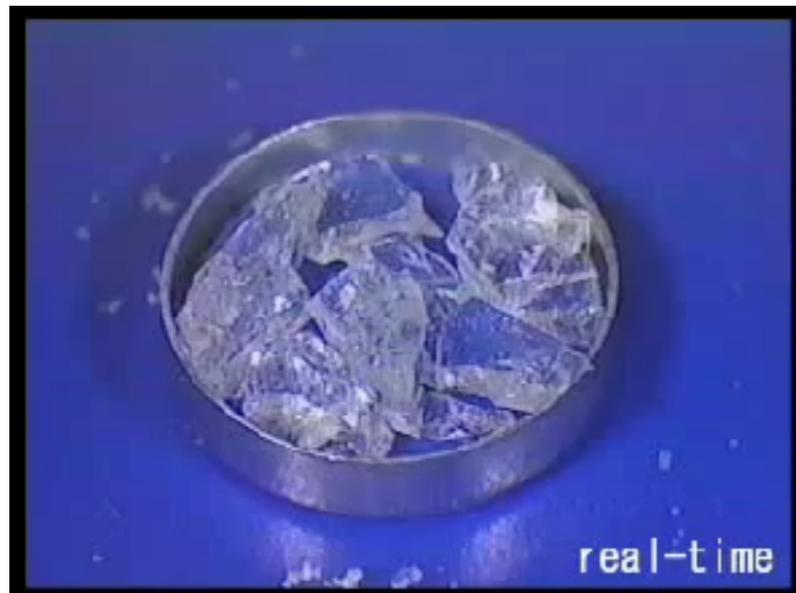
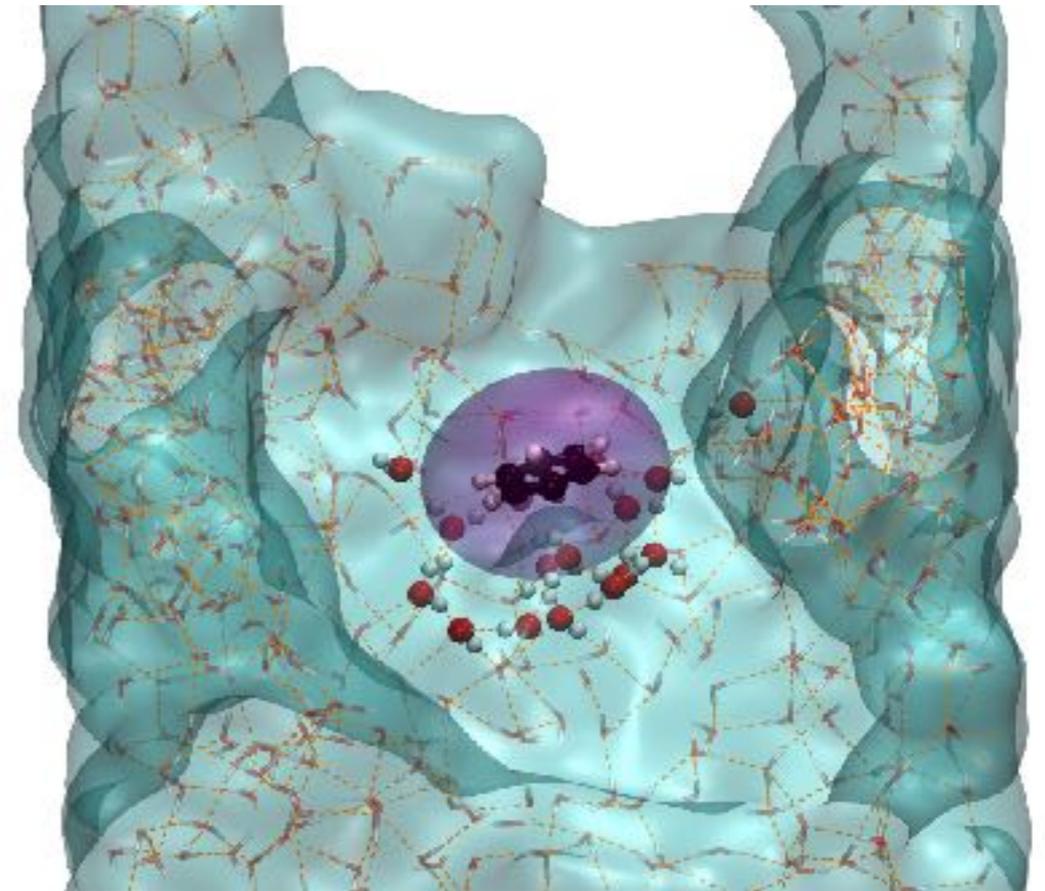




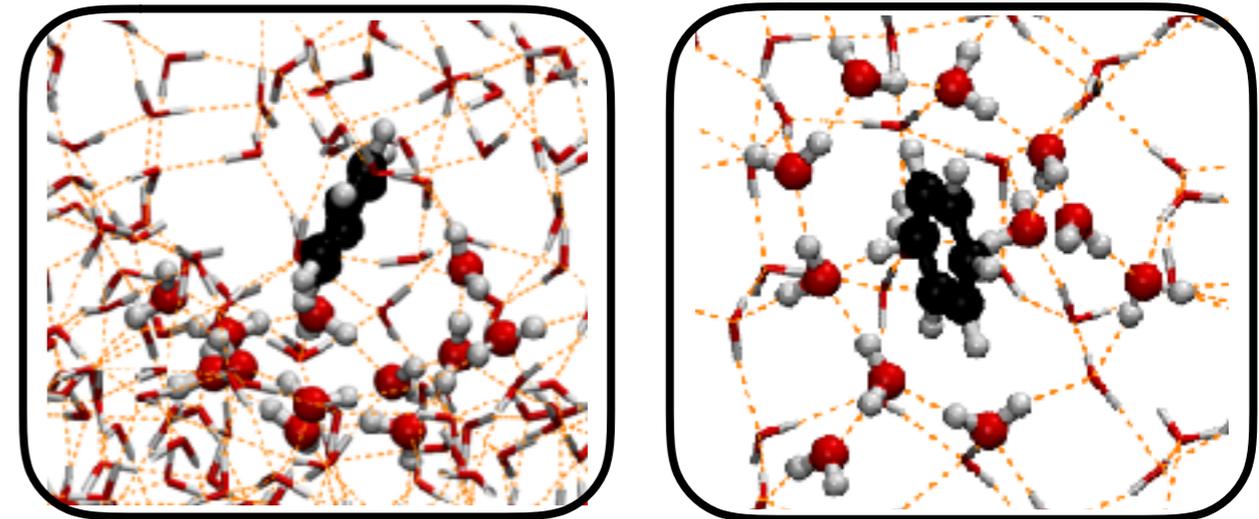
Sticky space ice?

Space ice model: Low density ice

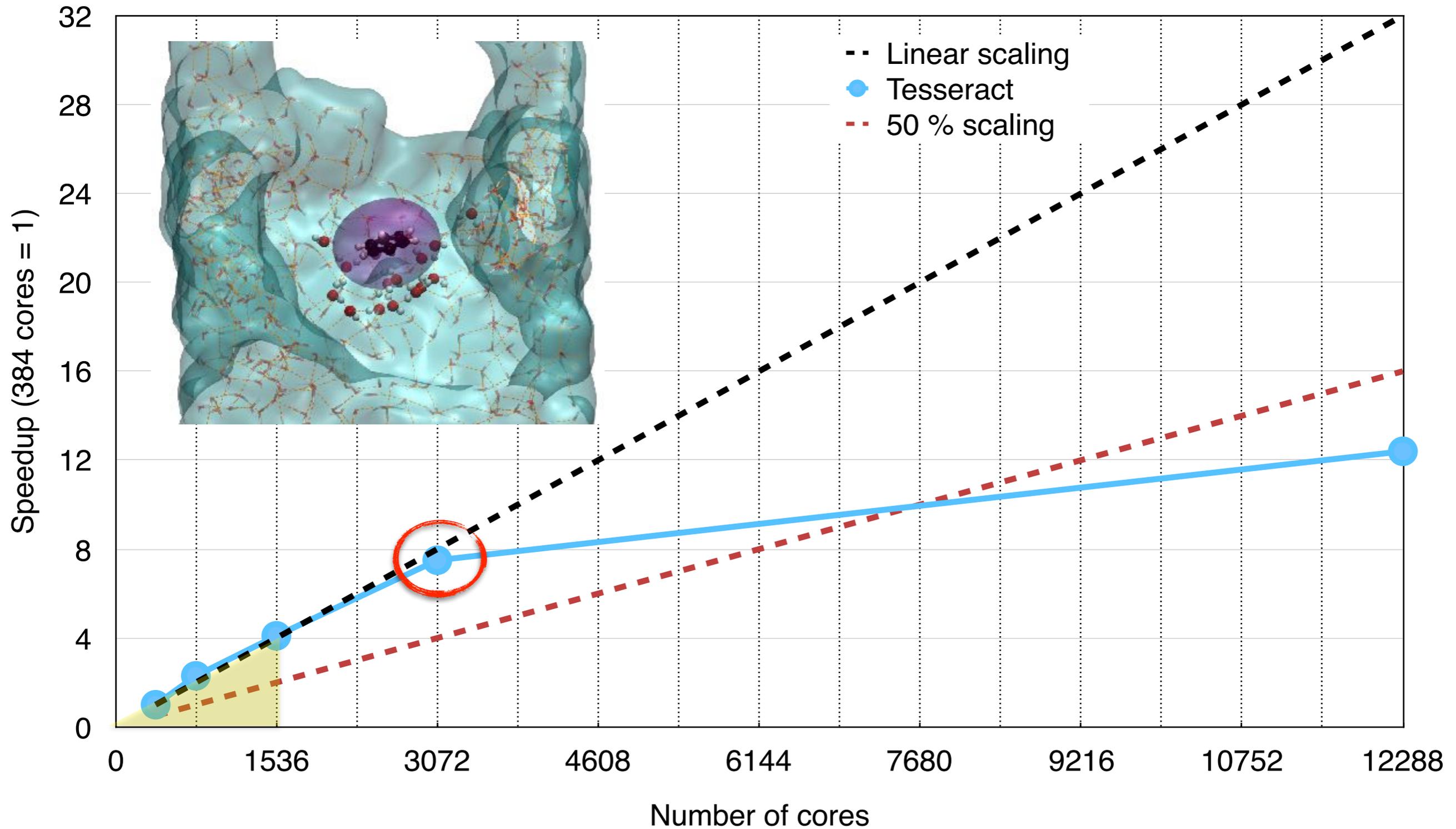
- Amorphous low density ice model with a single benzene molecule, 1512 atoms, 4030 electrons, PBE functional and a 200 Ry cutoff
- 10-fold size increase from C₁₂₀



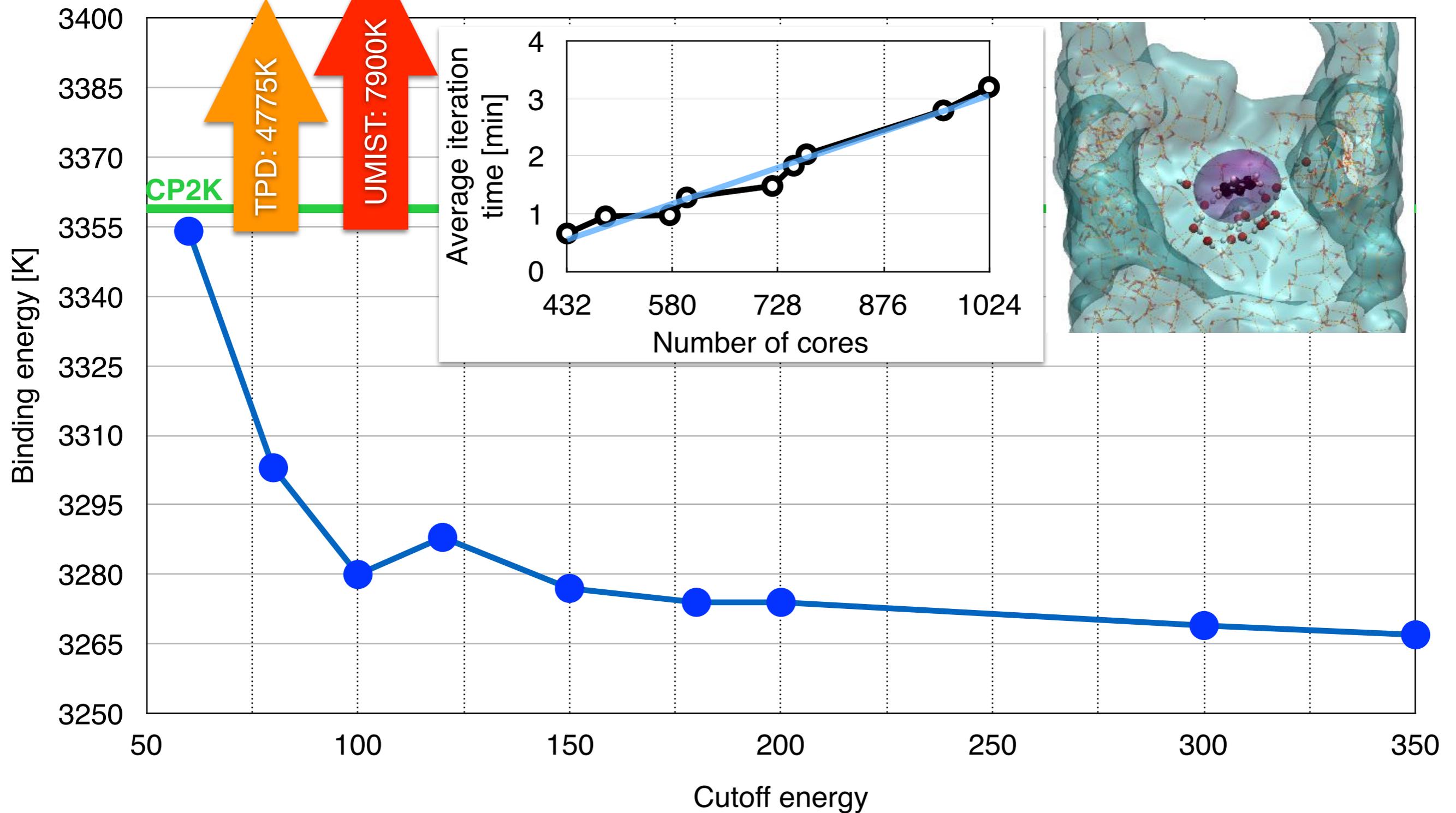
http://www.nims.go.jp/water/hda_lda_tr.html



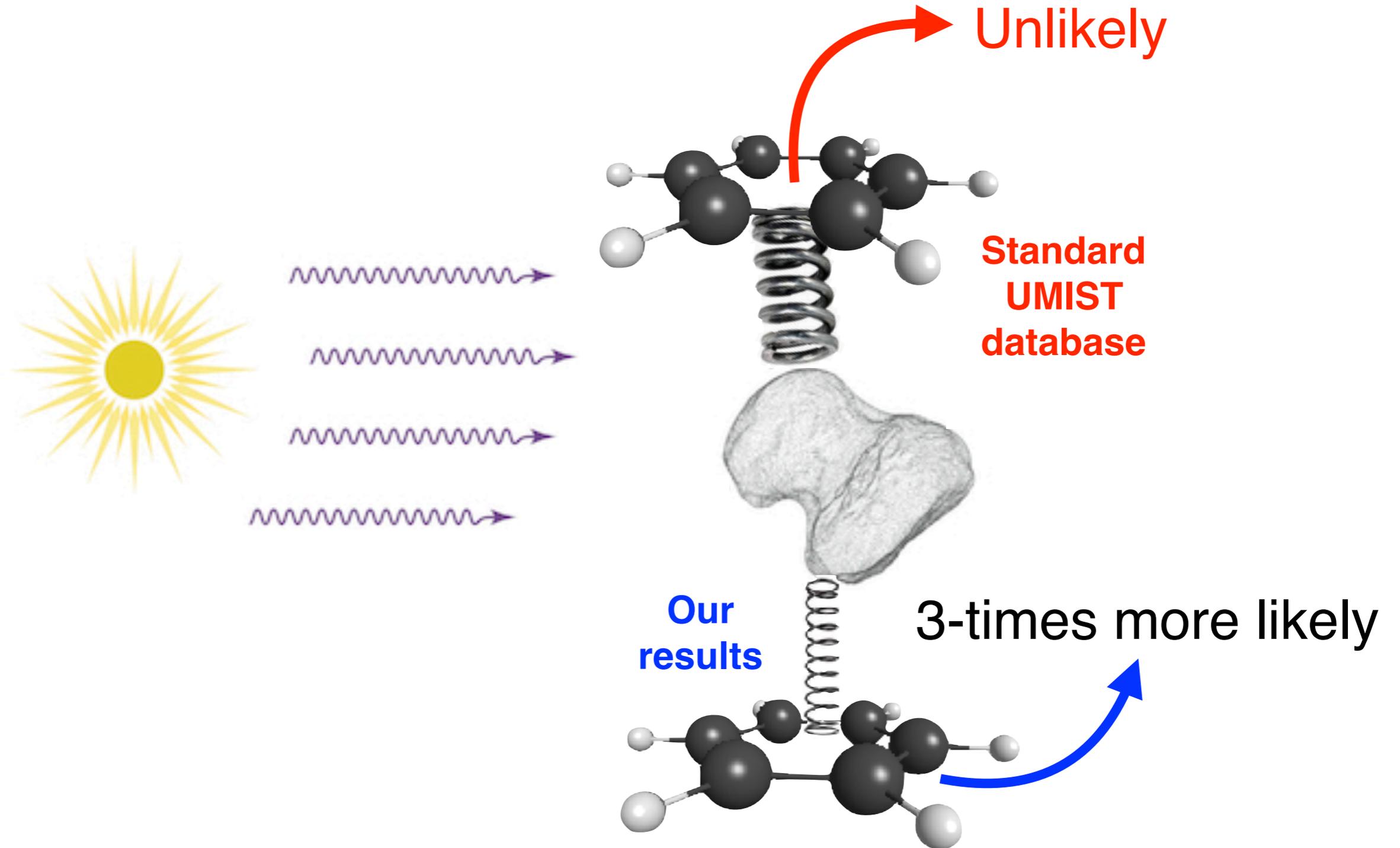
Tesseract scaling



Adsorption energy estimation



Why is it important?



Conclusions

- Omni-Path performance are similar to EDR in conventional scaling region but **better** in non-scalable regime
- Combination of enhanced locality, reduced MPI ranks and FFT grouping leads to predictable **scalability**
- However, understanding application and problem size are **key** to scaling at large core counts

Acknowledgements



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